

Mark Scheme (Results)

October 2023

Pearson Edexcel International Advanced Level In Physics (WPH15)

Paper 01: Thermodynamics, Radiation,

Oscillations and Cosmology

Question Number	Answer	Mark
1	A is the only correct answer	(1)
	B is not the correct answer, as temperature must be high for fusion	
	C is not the correct answer, as density must be high for fusion	
	D is not the correct answer, as temperature and density must be high for fusion	
2	A is the only correct answer	(1)
	B is not the correct answer, as parallax measurements do not involve intensity	
	C is not the correct answer, as parallax measurements do not involve luminosity	
	D is not the correct answer, as parallax measurements do not involve the Hubble	
	constant	
3	D is the only correct answer	(1)
	A is not the correct answer, as B.E./nucleon has a maximum for ⁵⁶ Fe	
	B is not the correct answer, as B.E./nucleon has a maximum for ⁵⁶ Fe	
	C is not the correct answer, as B.E./nucleon has a maximum for ⁵⁶ Fe	
4	B is the only correct answer	(1)
	A is not the correct answer, as acceleration is always towards the equilibrium point	
	C is not the correct answer, as acceleration is always towards the equilibrium point	
	D is not the correct answer, as this would increase the energy of oscillation	
5	D is the only correct answer	(1)
	A is not the correct answer, as motion does not change the wavelength of emission	
	B is not the correct answer, as motion does not change the wavelength of emission	
	C is not the correct answer, as the wavelength increases when the source is receding	
6	B is the only correct answer	(1)
	A is not the correct answer, as gravitational potential increases	
	C is not the correct answer, as gravitational force decreases and gravitational potential	
	increases	
	D is not the correct answer, as gravitational force decreases	
7	B is the only correct answer, as $F = mg$ and $g = (9.81 \text{ m s}^{-2})/4$	(1)
8	B is the only correct answer	(1)
	A is not the correct answer, as penetration is high	
	C is not the correct answer, as ionising power is low and penetration is high	
	D is not the correct answer, as ionising power is low	
9	B is the only correct answer	(1)
	A is not the correct answer, as main sequence stars to not go direct to white dwarfs	
	C is not the correct answer, as stars do not move down the main sequence	
	D is not the correct answer, as red giants do not return to the main sequence	
10	A is the only correct answer, as $T = 2\pi \sqrt{\frac{\ell}{g}}$	(1)

Question Number	Answer	Mark
11	Use of $pV = NkT$ to calculate T or kT (1)	
	Use of $\frac{1}{2}m\langle c^2\rangle = \frac{3}{2}kT$ (1)	
	[use of $\frac{1}{2}m\langle c^2\rangle = \frac{3pV}{2N}$ gets MP1 and MP2]	
	$\frac{1}{2}m\langle c^2\rangle = 5.9 \times 10^{-21} \text{J} \tag{1}$	3
	Example of calculation	
	$T = \frac{1.15 \times 10^5 \text{ Pa} \times 1.77 \times 10^{-3} \text{ m}^3}{5.15 \times 10^{22} \times 1.38 \times 10^{-23} \text{ J K}^{-1}} = 286 \text{ K}$	
	$\frac{1}{2}m\langle c^2\rangle = \frac{3}{2} \times 1.38 \times 10^{-23} \text{J K}^{-1} \times 286 \text{K} = 5.93 \times 10^{-21} \text{J}$	
	Total for question 11	3

Question Number	Answer	Mark
12	Two pairs of p , V readings from graph (1))
	Additional pair(s) of p , V readings from graph (1))
	$pV = 0.66 \ (\times \ 10^3 \ Pa \ m^3)$ [calculated for at least one pair of p , V readings]	
	Comment that value of pV is constant and so the student's claim is valid [dependent upon pV calculated for at least two pairs of p , V readings]	4
	Example of calculation	
	p = 110 kPa, V = 0.006 m $pV = 110 \times 10^3 \text{ Pa} \times 0.006 \text{ m}^3 = 660 \text{ Pa m}^3$	
	$p = 60 \text{ kPa}, V = 0.011 \text{ m}^3$ $pV = 60 \times 10^3 \text{ Pa} \times 0.011 \text{ m}^3 = 660 \text{ Pa m}^3$	
	$p = 51 \text{ kPa}, V = 0.013 \text{ m}^3$ $pV = 51 \times 10^3 \text{ Pa} \times 0.013 \text{ m}^3 = 663 \text{ Pa m}^3$	
	Total for question 12	4

Question Number	Answer		Mark
13(a)	Calculation of mass difference	(1)	
	Use of $\Delta E = c^2 \Delta m$	(1)	
	Conversion of energy from J to eV	(1)	
	E = 1.2 (MeV)	(1)	4
	Example of calculation		
	$(2.82185 \times 10^{-26} + 1.67299 \times 10^{-27}) - (2.32451 \times 10^{-26} + 6.64432 \times 10^{-27})$		
	$= (2.98915 - 2.98894) \times 10^{-26} = 2.07 \times 10^{-30} \text{ kg}$		
	$\Delta E = (3.0 \times 10^8 \text{ m s}^{-1})^2 \times 2.07 \times 10^{-30} \text{ kg} = 1.863 \times 10^{-13} \text{ J}$		
	$\Delta E = \frac{1.89 \times 10^{-13} \text{ J}}{1.6 \times 10^{-19} \text{ J eV}^{-1}} = 1.16 \times 10^6 \text{ eV} = 1.16 \text{ MeV}$		
13(b)	Momentum (and energy) is conserved	(1)	
	(So) products must have E_k / momentum after the reaction (as the alpha particle has momentum before the reaction)		
		(1)	2
	Total for question 13		6

Question Number	Answer		Mark
14(a)	The light/radiation (received) from the galaxies is red shifted Or Wavelength of light/radiation (received) from the galaxies was longer than expected	(1)	1
14(b)	EITHER		
	A straight line through the origin would be consistent with Hubble's expression	(1)	
	There is scatter about the line but the points are distributed evenly	(1)	
	So the expression may be valid (dependent upon MP2)	(1)	
	OR		
	A straight line through the origin would be consistent with Hubble's expression	(1)	
	(But) there are outliers and these are far from the line Or (But) only some of the points are close to the line	(1)	
	So the expression may not be valid (dependent upon MP2)	(1)	
	OR		
	The gradient of the line is equal to H_0	(1)	
	There is scatter about the line, so the value of H_0 is uncertain	(1)	
	So the expression may not be valid (dependent upon MP2)	(1)	3
	Total for question 14		4

Question Number			Answer				Mark
*15	structured ans Marks are awa and shows lin	wer with link arded for indi- es of reasoning table shows l	dent's ability to show ages and fully-sustair cative content and for g. how the marks should	ed reasoning how the an	ng. Iswer is stru	actured	
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	linkages and throughout Answer is pa of reasoning	fully sustaine	ed lines of reasoning of red with some linkag	demonstrates and line	1		
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	3	3 2	1	3			
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	therm Or St	istor is at the op heating an	nometer near to the the temperature measured wait before taking	ed by the the readings	nermomete	r)	
	Or Sv	vitch current	ent/p.d. (to prevent it off between reading eter at eye level	_	e thermisto	r)	
	Total for que		y • •				6

Question Number	Answer	Mark
16(a)	Use of $\rho = \frac{m}{V}$ (1)	
	Use of $\Delta E = mc\Delta\theta$ (1)	
	Use of $P = \frac{\Delta E}{\Delta t}$ (1)	
	P = 1630 (W) (1)	4
	Example of calculation	
	$m = 4.25 \times 10^{-4} \text{ m}^3 \times 998 \text{ kg m}^{-3} = 0.424 \text{ kg}$	
	$\Delta E = 0.424 \text{ kg} \times 4190 \text{ J kg}^{-1} \text{K}^{-1} \times (100 - 22) \text{ K} = 1.386 \times 10^5 \text{ J}$	
	$P = \frac{1.386 \times 10^5 \mathrm{J}}{85 \mathrm{s}} = 1631 \mathrm{W}$	
16(b)	Use of $\Delta E = L\Delta m$ (1)	
	Use of $P = \frac{\Delta E}{\Delta t}$ (1)	
	$t = 440 \text{ s (ecf from (a))} \tag{1}$	3
	Example of calculation	
	$\Delta E = 0.75 \times 0.424 \text{ kg} \times 2.26 \times 10^6 \text{ J kg}^{-1} = 7.19 \times 10^5 \text{ J}$	
	$t = \frac{7.19 \times 10^5 \mathrm{J}}{1630 \mathrm{W}} = 441 \mathrm{s}$	
	Total for question 16	7

Question Number	Answer		Mark
17(a)	Use of $g = \frac{GM}{r^2}$	(1)	
	$g = 0.40 \text{ N kg}^{-1}$ Example of calculation	(1)	2
	$\frac{EXAMPLE of Calculation}{g = \frac{6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2} \times 3.1 \times 10^{21} \text{ kg}}{(7.15 \times 10^5 \text{ m})^2} = 0.404 \text{ N kg}^{-1}$ Equates $F = \frac{GMm}{r^2}$ with $F = m\omega^2 r$		
17(b)	Equates $F = \frac{GMm}{r^2}$ with $F = m\omega^2 r$	(1)	
	Use of $\omega = \frac{2\pi}{T}$	(1)	
	$T_{\rm M} = 9.7 \times 10^9 { m s}$	(1)	
	Conversion between seconds and years [Must see a unit for <i>T</i> , either in MP3 or MP4]	(1)	
	Calculates ratio of orbital time of Makemake with orbital time of Pluto [Ratio includes a percentage calculation]	(1)	
	Comparison of values and consistent conclusion	(1)	
	OR	(4)	
	Equates $F = \frac{GMm}{r^2}$ with $F = \frac{mv^2}{r}$	(1)	
	Use of $v = \frac{2\pi r}{T}$	(1) (1)	
	$T_{\rm M}=9.7\times10^9~{\rm s}$	(1)	
	Conversion between seconds and years	(1)	
	Calculates ratio of orbital time of Makemake with orbital time of Pluto [Ratio includes a percentage calculation]	(1)	
	Comparison of values and consistent conclusion	(1)	6
	Example of calculation		
	$\frac{GMm}{r^2} = m\omega^2 r$		
	$\omega = \sqrt{\frac{GM}{r^3}} = \sqrt{\frac{6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-1} \times 1.99 \times 10^{30} \text{ kg}}{(6.80 \times 10^{12} \text{ m})^3}}$		
	$\omega = 6.50 \times 10^{-10} \text{ rad } s^{-1}$		
	$T = \frac{2\pi}{\omega} = \frac{2\pi \text{ rad}}{6.50 \times 10^{-10} \text{ rad s}^{-1}} = 9.67 \times 10^9 \text{ s} = \frac{9.67 \times 10^9 \text{ s}}{3.15 \times 10^7 \text{ s year}^{-1}}$ $= 307 \text{ year}$		
	orbital time ratio = $\frac{307 \text{ year}}{248 \text{ year}} = 1.24$		
	The orbital time of Makemake is 24% greater than that of Pluto, so website statement is not quite accurate		
	Total for question 17		8

Question Number	Answer		Mark
18(a)	Use of $V = \frac{4}{3}\pi r^3$	(1)	
	Use of $\rho = \frac{m}{V}$	(1)	
	Use of $F = \frac{Gm_1m_2}{r^2}$	(1)	
	$F = 7.4 \times 10^5 \mathrm{N}$	(1)	4
	$\frac{\text{Example of calculation}}{V = \frac{4}{3}\pi r^3 = \frac{4}{3}\pi \left(\frac{5.65 \text{ m}}{2}\right)^3 = 94.437 \text{ m}^3}$ $m = \rho V = 1950 \text{ kg m}^{-3} \times 94.437 \text{ m}^3 = 1.842 \times 10^5 \text{ kg}$ $F = \frac{Gm_1m_2}{r^2}$ $= \frac{6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-3} \times 5.98 \times 10^{24} \text{ kg} \times 1.842 \times 10^5 \text{ kg}}{(6.38 \times 10^6 \text{ m} + 3.59 \times 10^6 \text{ m})^2}$		
	$\therefore F = 7.39 \times 10^5 \text{ N}$		
18(b)	$\therefore F = 7.39 \times 10^5 \text{ N}$ Use of $V_{\text{grav}} = (-) \frac{GM}{r}$	(1)	
	Use of $E_{\text{grav}} = m \times V_{\text{grav}}$	(1)	
	$\therefore \Delta E_{\text{grav}} = (-) \text{ 4.1} \times 10^{12} \text{ J (Allow ecf for mass from (a))}$	(1)	3
	[Either mass can be used for M in the potential equation, but to award MP2 the multiplier m . must not be the mass used in the potential equation.]		
	$\frac{\text{Example of calculation}}{\Delta E_{\text{grav}} = -6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2} \times 1.842 \times 10^5 \text{ kg} \times 5.98 \times 10^{-24} \text{ kg}}{\times \left(\frac{1}{6.38 \times 10^6 \text{ m}} - \frac{1}{(6.38 \times 10^6 + 3.59 \times 10^6) \text{ m}}\right)}$		
	$\therefore \Delta E_{\rm grav} = -4.14 \times 10^{12} \mathrm{J}$		
18(c)	Work would be done on the asteroid by frictional forces Or Drag/friction causes heating (of the asteroid)	(1)	
	Asteroid burns up	(1)	2
	Total for question 18		9

Question Number	Answer		Mark
19(a)(i)	Use of $\lambda = \frac{\ln 2}{t_{1/4}}$	(1)	
	$\lambda = 7.31 \times 10^{-10} (s^{-1})$	(1)	2
	$\lambda = \frac{\ln 2}{30.1 \times 3.15 \times 10^7 \text{ s}} = 7.31 \times 10^{-10} \text{ s}^{-1}$		
19(a)(ii)	Use of $\frac{dN}{dt} = -\lambda N$	(1)	
	Use of $u = 1.66 \times 10^{-27}$ kg with 137	(1)	
	$m = 5.9 \times 10^{-6} \text{ (kg) (Allow ecf from (a)(i))}$	(1)	3
	$N = \frac{19 \times 10^9 \text{ s}^{-1}}{7.31 \times 10^{-10} \text{ s}^{-1}} = 2.60 \times 10^{19}$		
	$m = 2.60 \times 10^{19} \times 137 \times 1.66 \times 10^{-27} \text{kg} = 5.91 \times 10^{-6} \text{kg}$		
19(a)(iii)	Use of $A = A_0 e^{-\lambda t}$	(1)	
	A = 18.1 GBq (Allow ecf from (a)(i))	(1)	2
	Example of calculation		
	$A = 19 \times 10^{9} \text{Bq} \times \text{e}^{-7.31 \times 10^{-10} \text{s}^{-1} \times 2 \times 3.15 \times 10^{7} \text{s}}$		
	$A = 1.81 \times 10^{10} \text{ Bq}$		
19(b)	Use of total energy released = $\left(\frac{\Delta N}{\Delta t}\right) \times \Delta t \times E$		
	Or Use of total energy released $\stackrel{\triangle L}{=} \Delta N \times E$	(1)	
	Use of 1 eV = 1.6×10^{-19} J	(1)	
	Total energy released = 4.3×10^3 (J)	(1)	3
	[If $\left(\frac{\Delta N}{\Delta t}\right) \times \Delta t$ determined by using exponential decay equation to calculate number of undecayed nuclei after 14 days; final answer should round to 4300 (J)]		
	Example of calculation $E = 19 \times 10^9 \text{ s}^{-1} \times 14 \times 86400 \text{ s} \times 1.17 \text{ MeV} = 2.69 \times 10^{16} \text{ MeV}$ $E = 2.69 \times 10^{16} \text{ MeV} \times 10^6 \times 1.6 \times 10^{-19} \text{ J eV}^{-1} = 4.30 \times 10^3 \text{ J}$		
	Total for question 19		10

Question Number	Answer		Mark
20(a)	There is a (resultant) force that is proportional to the displacement from the equilibrium position	(1)	
	and (always) acting towards the equilibrium position	(1)	2
		(1)	2
	[Allow references to acceleration; an equation with symbols defined correctly is a valid response for both marks.]		
20(b)	EITHER	(1)	
	Use of $F = mg$	(1)	
	Use of $\Delta F = (-)k\Delta x$	(1)	
	Use of $T = 2\pi \sqrt{\frac{m}{k}}$	(1)	
	Use of $\omega = \frac{2\pi}{T}$	(1)	
	Use of $v = \omega x_0 \sin \omega t$	(1)	
	$v_{\rm max} = 0.34 \mathrm{ms^{-1}}$	(1)	
	OR	()	
	Use of $F = mg$	(1)	
	Use of $\Delta F = (-)k\Delta x$	(1)	
	Use of $\Delta E_{el} = \frac{1}{2} F \Delta x$	(1)	
	Use of $E_k = \frac{1}{2}mv^2$	(1)	
	Use of energy conservation	(1)	
	$v_{\rm max} = 0.34 \; {\rm m s^{-1}}$	(1)	6
	Example of calculation $F = 0.150 \text{ kg} \times 9.81 \text{ N kg}^{-1} = 1.47 \text{ N}$ $k = \frac{1.47 \text{ N}}{7.5 \times 10^{-2} \text{ m}} = 19.6 \text{ N m}^{-1}$		
	$T = 2\pi \sqrt{\frac{0.150 \text{ kg}}{19.6 \text{ N m}^{-1}}} = 0.549 \text{ s}$		
	$\omega = \frac{2\pi \text{rad}}{0.549 \text{s}} = 11.4 \text{rad s}^{-1}$		
	$v_{\text{max}} = 11.4 \text{ rad s}^{-1} \times 3.0 \times 10^{-2} \text{ m} = 0.343 \text{ m s}^{-1}$		
20(c)	Energy is transferred out of the oscillating system		
	Or energy is dissipated (to surroundings)	(1)	
	Because work is done by/against resistive forces	(1)	2
	[Allow MAX 1 for reference to damping]		10
	Total for question 20		10

Question Number	Answer	Mark
21(a)(i)	Use of $\lambda_{max}T = 2.898 \times 10^{-3}$ (1)	
	Use of $L = \sigma A T^4$ and $A = 4\pi r^2$ Or Use of $L = \sigma A T^4$ to calculate A and $A \propto r^2$ $\frac{r_B}{r_S} = 990$ (1)	
	$\frac{r_B}{r_S}$ is approximately equal to 1000, so claim is accurate Or $\frac{r_B}{r_S}$ is less than 1000, so claim is inaccurate Or $\frac{r_B}{r_S}$ is not equal to 1000, so claim is inaccurate (Allow use of calculated ratio with consistent conclusion)	4
	$\frac{\text{Example of calculation}}{T = \frac{2.898 \times 10^{-3} \text{ m K}}{850 \times 10^{-9} \text{ m}} = 3410 \text{ K}$ $\frac{L_B}{L_S} = \frac{4\pi\sigma r_B^2 T_B^4}{4\pi\sigma r_S^2 T_S^4}$ $\frac{r_B}{r_S} = \sqrt{\frac{L_B}{L_S} \times \frac{T_S^4}{T_B^4}} = \sqrt{\frac{4.49 \times 10^{31} \text{ W}}{3.83 \times 10^{26} \text{ W}} \times \left(\frac{5800 \text{ K}}{3410 \text{ K}}\right)^4} = 991$	
21(a)(ii)	Sun in correct position (1) Betelgeuse in correct position (1) $L/L_{\text{Sun}} = 10^6$ 10^2 10^{-2} 10^{-4} $40000 = 20000 = 10000 = 5000 = 2500$	2
21(a)(iii)	T/K A main sequence star is a star that is fusing <u>hydrogen</u> in its <u>core</u> (1)	1
21(a)(iii)	A main sequence star is a star that is fusing <u>hydrogen</u> in its <u>core</u> (1)	1

21(b)	Use of $\omega = \frac{2\pi}{T}$ (1) Use of $v = r\omega$ (1) Use of $\frac{\Delta\lambda}{\lambda} = \frac{v}{c}$ (1) Determines range by taking 91.2 nm ± $\Delta\lambda$ (1) Maximum wavelength = 91.8 (nm) Minimum wavelength = 90.6 (nm) (1) $\frac{2\pi}{T} = \frac{2\pi \text{ rad}}{33.5 \times 10^{-3} \text{s}} = 187.6 \text{ rad s}^{-1}$ $v = 10.25 \times 10^3 \text{ m} \times 187.6 \text{ rad s}^{-1} = 1.922 \times 10^6 \text{ m s}^{-1}$ $\frac{\Delta\lambda}{91.2 \times 10^{-9} \text{ m}} = \frac{1.922 \times 10^6 \text{ m s}^{-1}}{3.00 \times 10^8 \text{ m s}^{-1}}$ $\therefore \Delta\lambda = 6.408 \times 10^{-3} \times 91.2 \times 10^{-9} \text{ m} = 5.84 \times 10^{-10} \text{ m}$	6
	Total for question 21	13